

Charmonium production off nuclei: from SPS to RHIC

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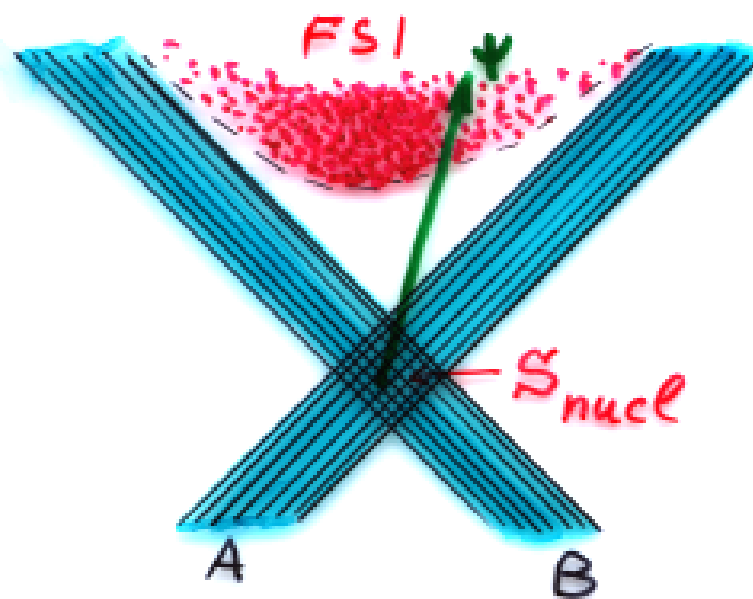
Charmonium production off nuclei: from SPS to RHIC

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Charmonium suppression as a probe for QGP



$$S_{AB}^{\psi} = S_{\text{nucl}} \times S_{\text{FSI}}$$

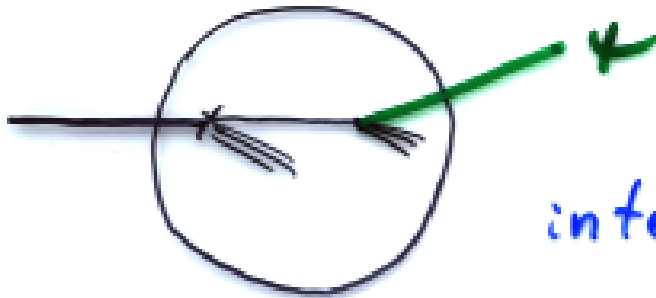
The base line The probe

S_{nucl} is still the main source of suppression, one desperately needs to understand it.

First of all, in pA collisions

Nuclear effects at SPS

1. Absorption



Initial state interaction is assumed to have no effect

$$S_{PA}^{\psi} = \frac{G_{PA}^{\psi}}{A G_{PN}^{\psi}} = \int d^2b \int_{-\infty}^{\infty} dz \rho_A(b, z) e^{-G_{abs}^{\psi N} \int_z^{\infty} dz' \rho_A(b, z')}$$

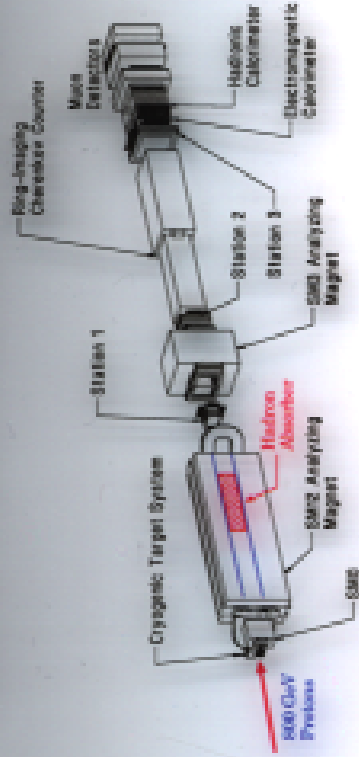
$$S_{PA}^{\psi} < 1$$

Does this simple model which misses many important effects, indeed describes data for ψ suppression in pA collisions?

Certainly not!

No x_F -dependence is predicted in a strict contradiction with data.

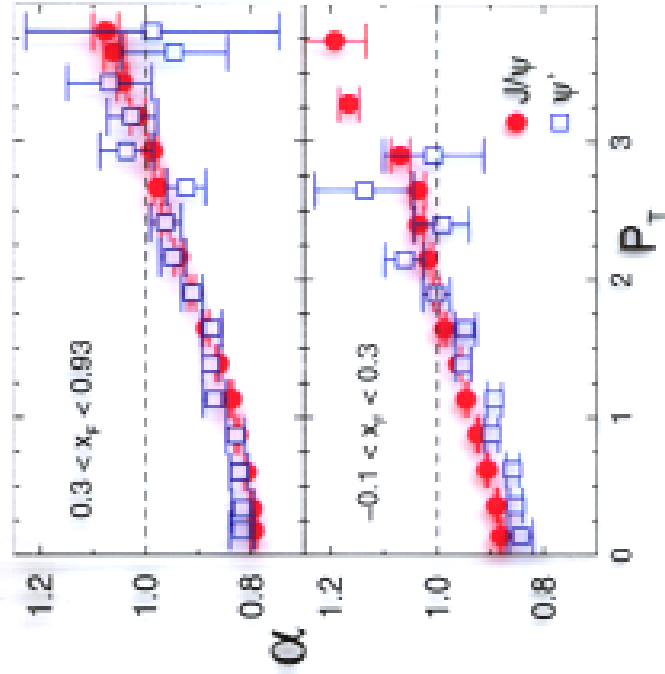
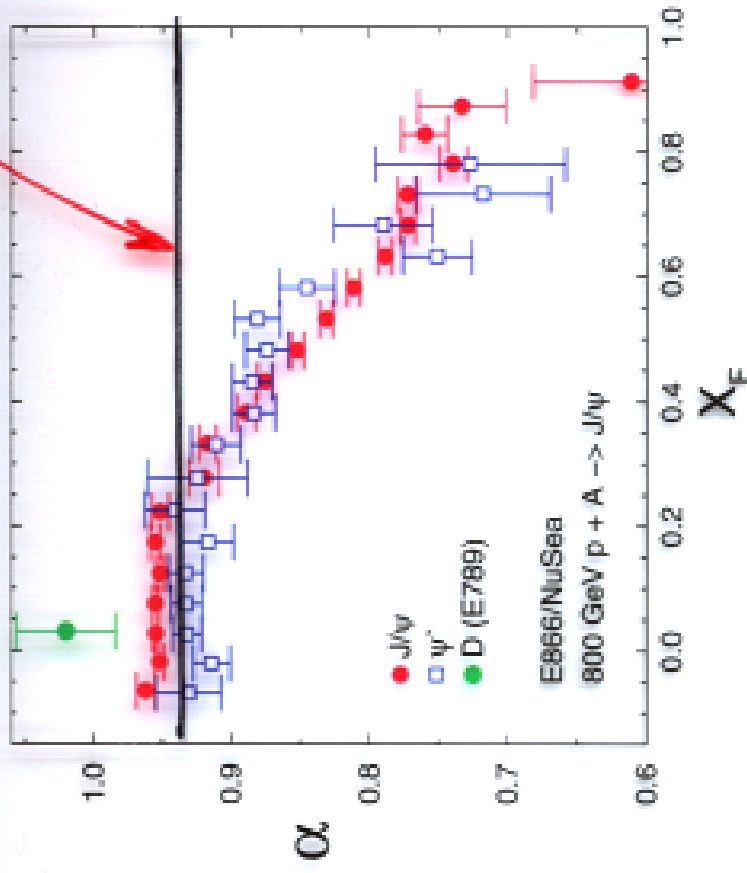
Nuclear Dependence of J/ψ and ψ' Production in 800 GeV/c p-A Collisions FNAL E866/NuSea



AGU, ANL, FNAL, GSU, IIT, LANL, LSU,
NMSU, UNM, ORNL, TAMU, Valpo.

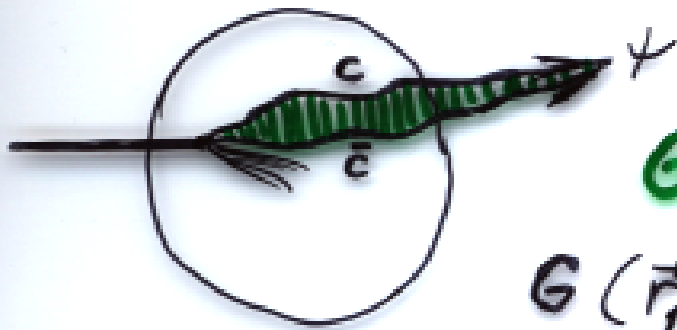
$$\sigma_A = \sigma_p * A^\alpha$$

absorption model



2. Formation time effects

B. K.
B. Zakharov
PRD44(1991) 3466



Green function

$$G(\vec{r}_1, z_1; \vec{r}_2, z_2)$$

Describes the propagation of a $c\bar{c}$ pair with initial separation \vec{r}_1 at the longitudinal coordinate z_1 , up to the final \vec{r}_2 at z_2 .

At fixed energy of Ψ the absorption cross section (effective) varies between

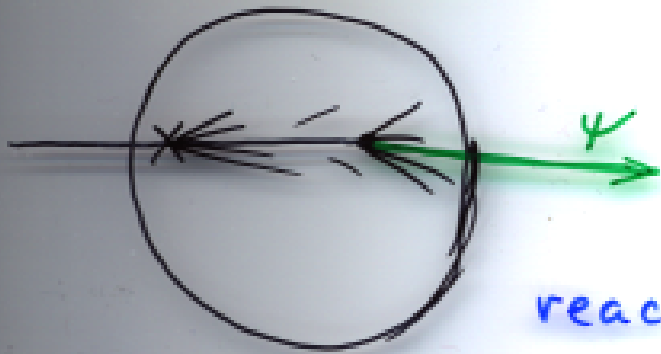
$$\sigma_{\text{eff}} = \frac{\langle \sigma_{q\bar{q}}^2(r) \rangle}{\langle \sigma_{q\bar{q}}(r) \rangle} \approx 3.5 \text{ mb}$$

for light nuclei, up to

$$\sigma_{\text{eff}} \approx 5.8 \text{ mb}$$

for very large nuclei (corrected for Ψ' and χ decays).

3. Energy loss



Similar to Drell-Yan reaction, initial state interactions lead to energy loss by the incident partons preceding production of ψ . It results in an extra suppression of the ψ production rate, especially at large x_F .

The recent analysis of the E772/866 data gives the rate of energy loss

$$\star - \frac{dE}{dz} = 2.3 \pm 0.52 \pm 0.5 \frac{\text{GeV}}{\text{fm}} \quad \text{M. Johnson et al.}$$

hep-ex/0010051

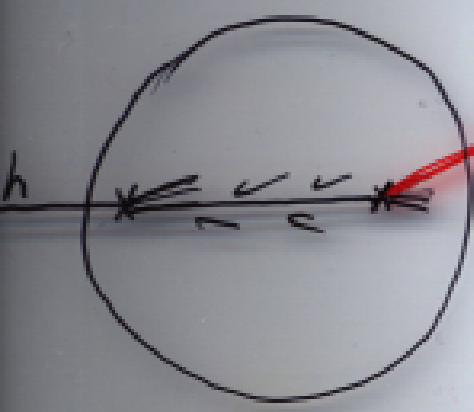
Manifestations of energy loss

25
4

- Charmonium (Drell-Yan pair) production off nuclei

The projectile hadron interacts inelastically at the nuclear surface and this is a start for hadronization.

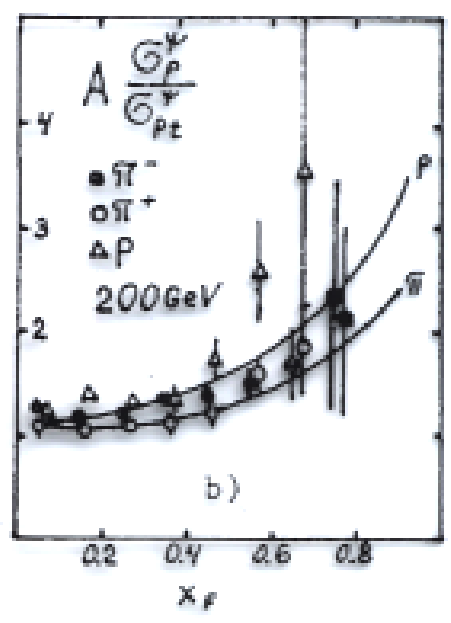
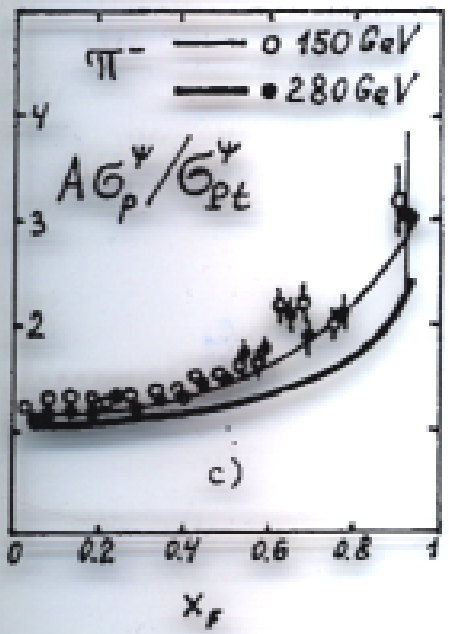
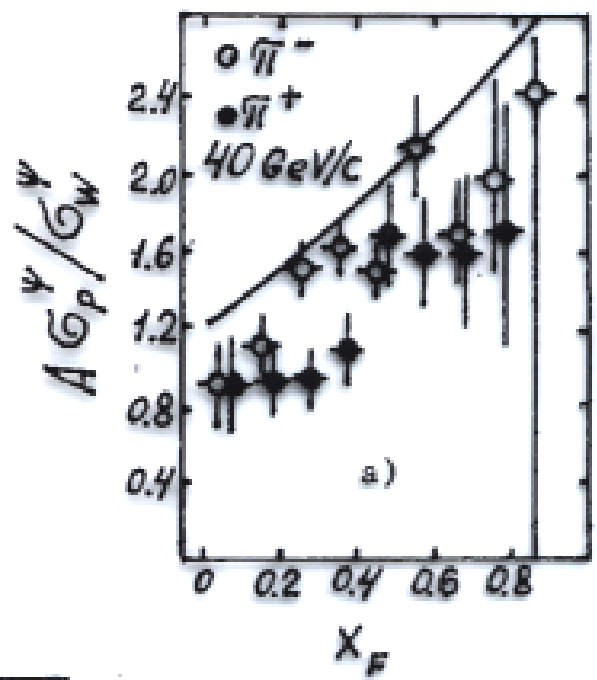
B.K. & F. Niedermayer
1984



$$\frac{dE}{dz} = -3 \text{ GeV/fm}$$

$$\tilde{E} = E - \alpha \Delta Z$$

$$\Delta X_1 = - \frac{\alpha \Delta Z}{E}$$

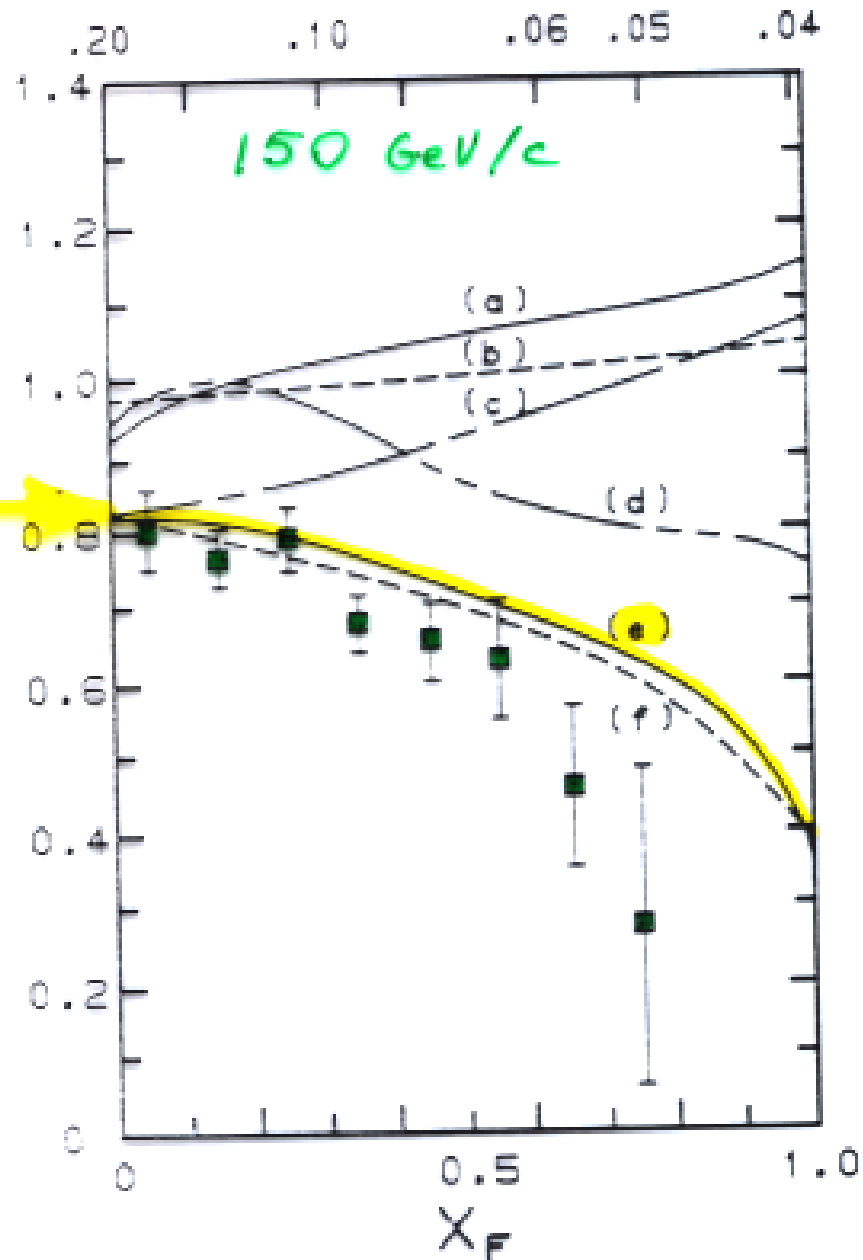


$\alpha = 1 \pm 1 \frac{\text{GeV}}{\text{fm}}$

E537

S. Katsanevas *et al.* X_2

PRL 60(1988)2121



B.K.

F. Niedermayer
1984

R1

 X_F

FIG. 5. Comparison of the ratio $(A^{-1}d\sigma/dx_F)$ for $\pi^- W$ to $\pi^- Be$ to various models: curve *a*, soft- π model; curve *b*, rescaling model; curve *c*, six-quark model; curve *d*, shadowing model; curve *e*, rescattering model; and curve *f*, three-gluon-fusion model without $q\bar{q}$ contributions.

4. Gluon antishadowing at large x_2

$$\frac{\sigma_{PA}^\psi}{\sigma_{PN}^\psi} \propto \frac{G_A(x_2)}{G_N(x_2)}$$

$$x_2 \approx \frac{M_\psi^2}{5x_1} = \frac{M_\psi^2}{2m_N E_\psi} \approx 0.1$$

Gluons are expected to be enhanced
by 10-20% in heavy nuclei

K. Eskola
 V. Kolhinen
 P. Ruuskanen

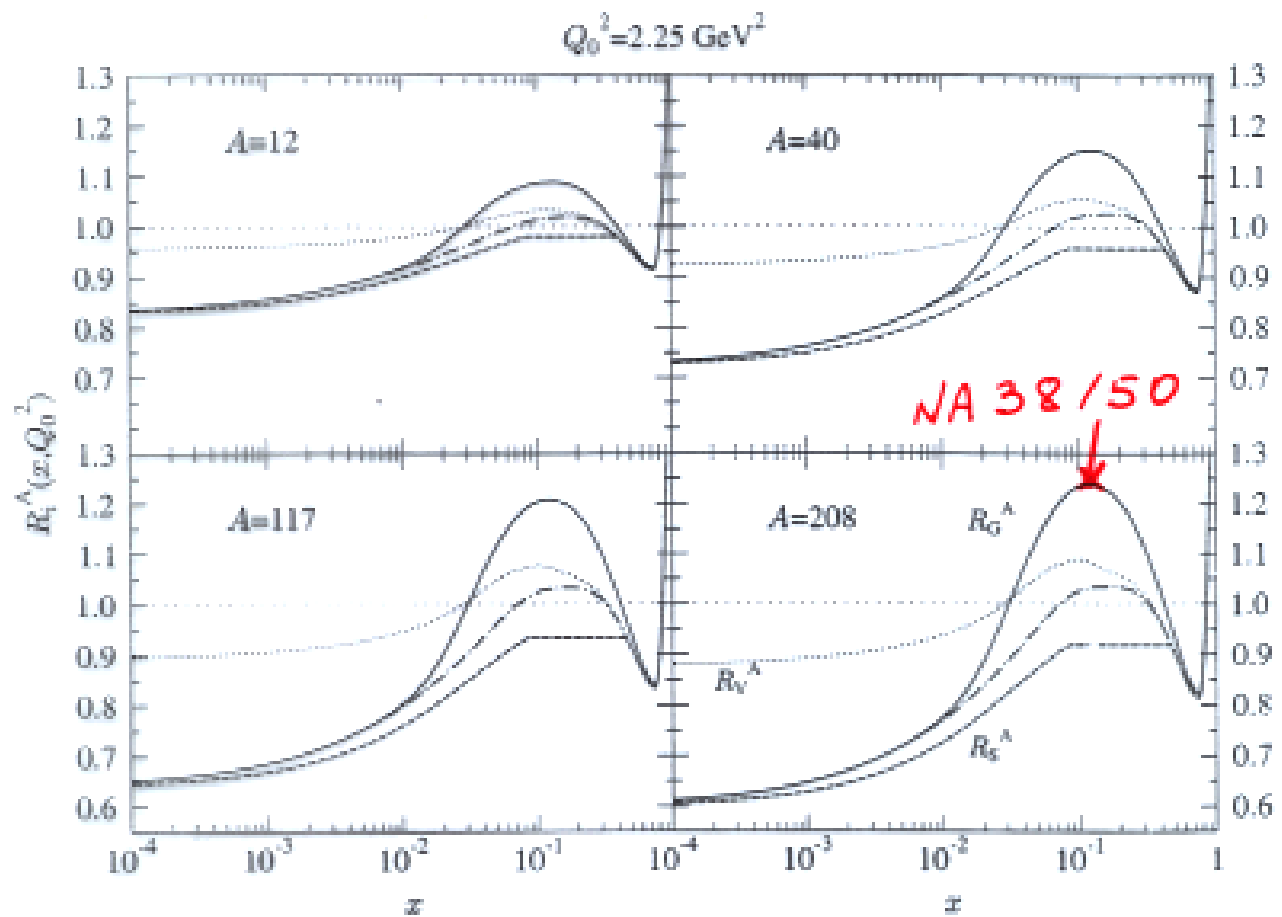


Figure 2: The initial nuclear ratios $R_G^A(x, Q_0^2)$ (solid line), $R_D^A(x, Q_0^2)$ (dotted) and $R_S^A(x, Q_0^2)$ (dashed) for isoscalar nuclei at $Q_0^2 = 2.25 \text{ GeV}^2$. The ratio $R_V^A(x, Q_0^2)$ (dotted-dashed) is also shown.

NA 38/50

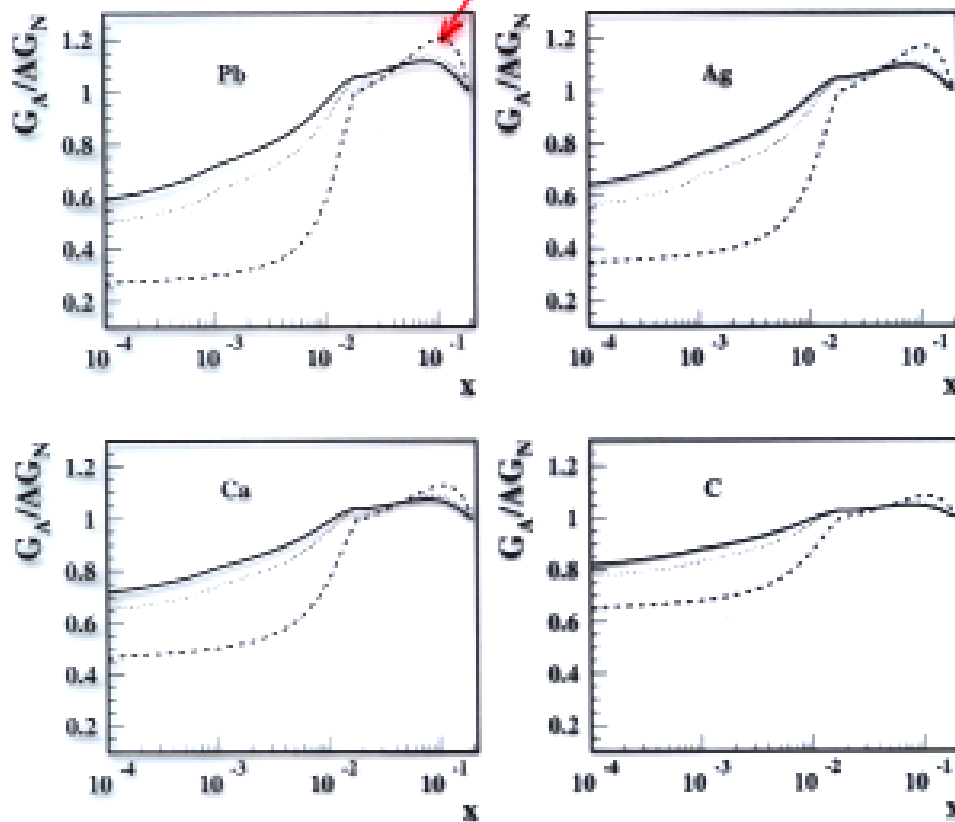
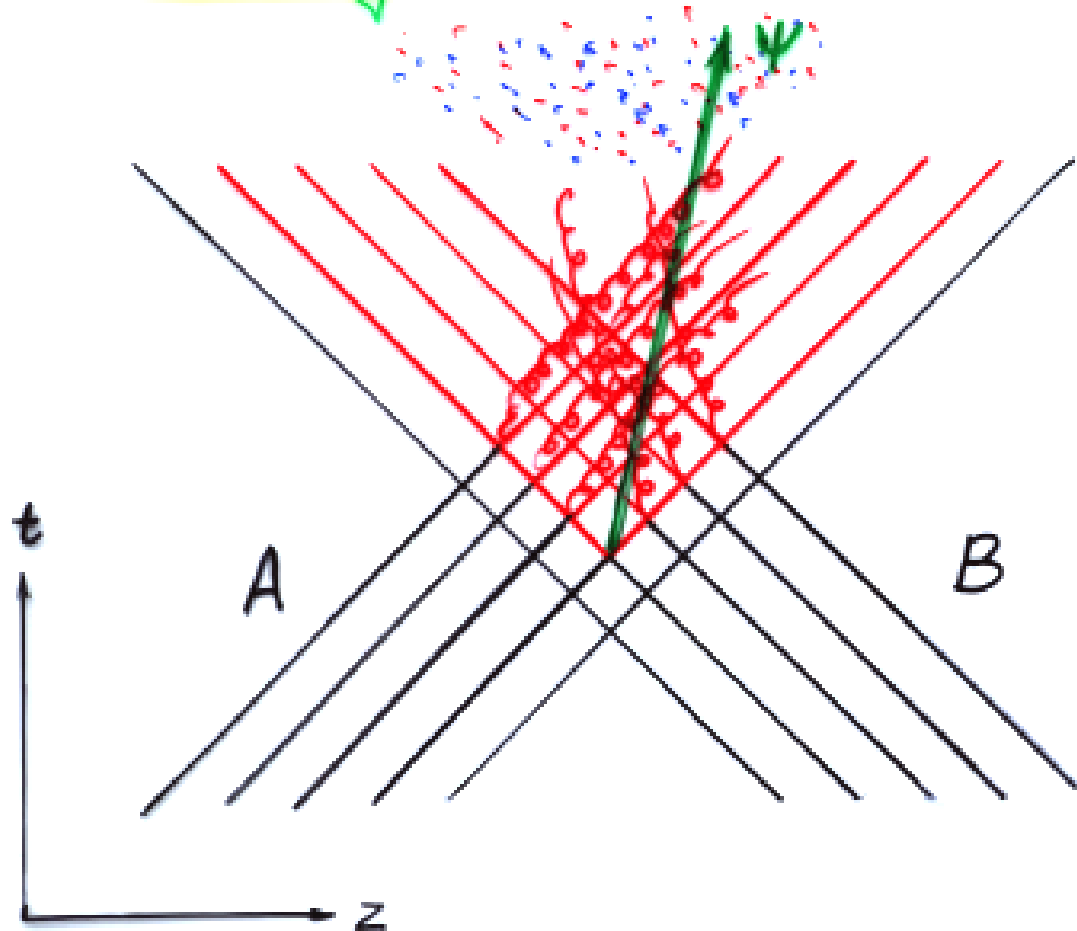


FIG. 4. Dependence of G_A/AG_N on x for $Q=2,5,10$ GeV (dashed, dotted, solid curves) calculated in the quasieikonal model.



5. Excitation of nuclear matter in heavy ion collisions



J. Hüfner
& B.K.

J. Hüfner
B.K.
& A. Polleri

Each nucleon crossing the Ψ trajectory has already interacted and is colored.

In addition the radiated gluons (those who lost coherence with the source) also contribute to the break up of the Ψ .

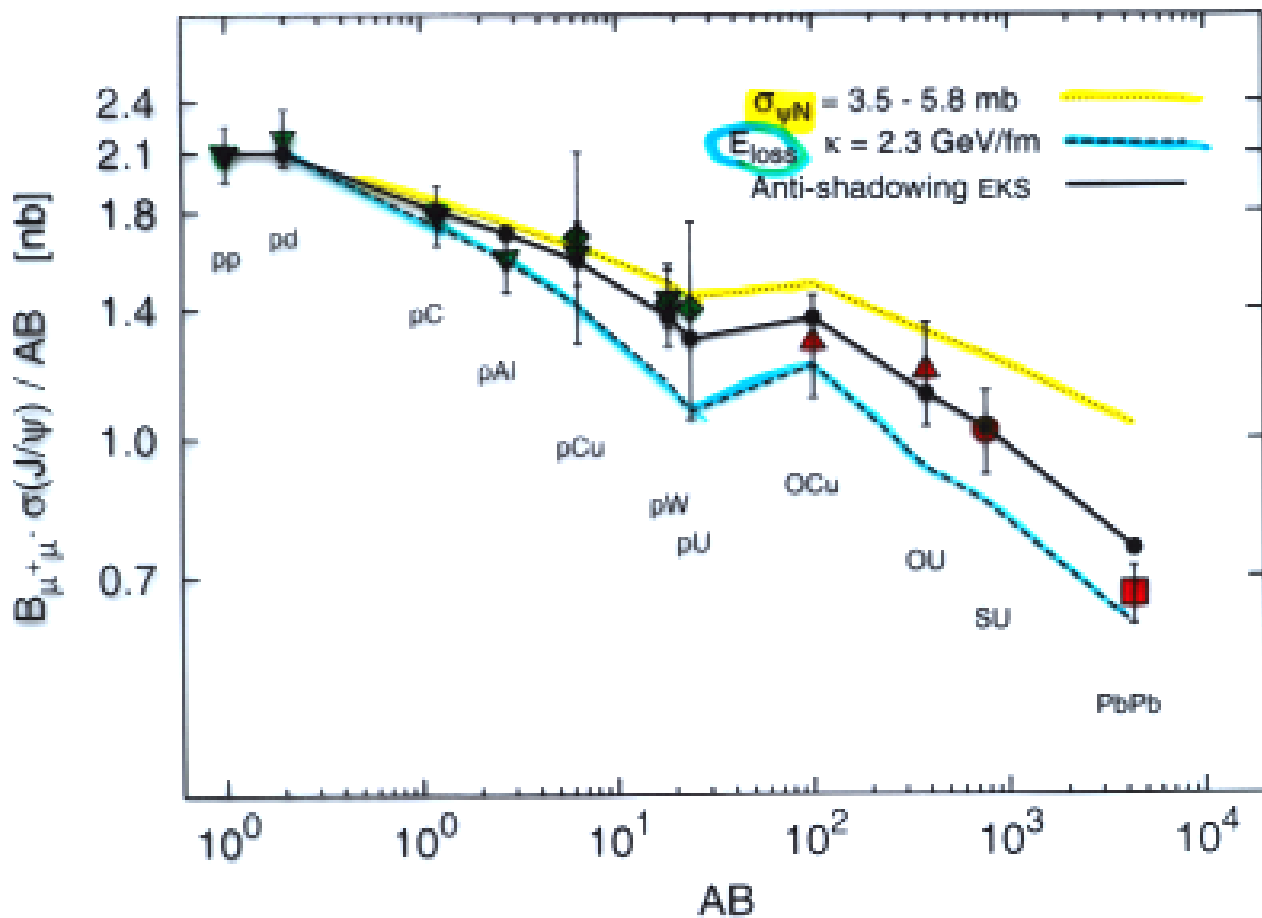
This is a very nonlinear effect.

It is practically absent for pA or light nuclei

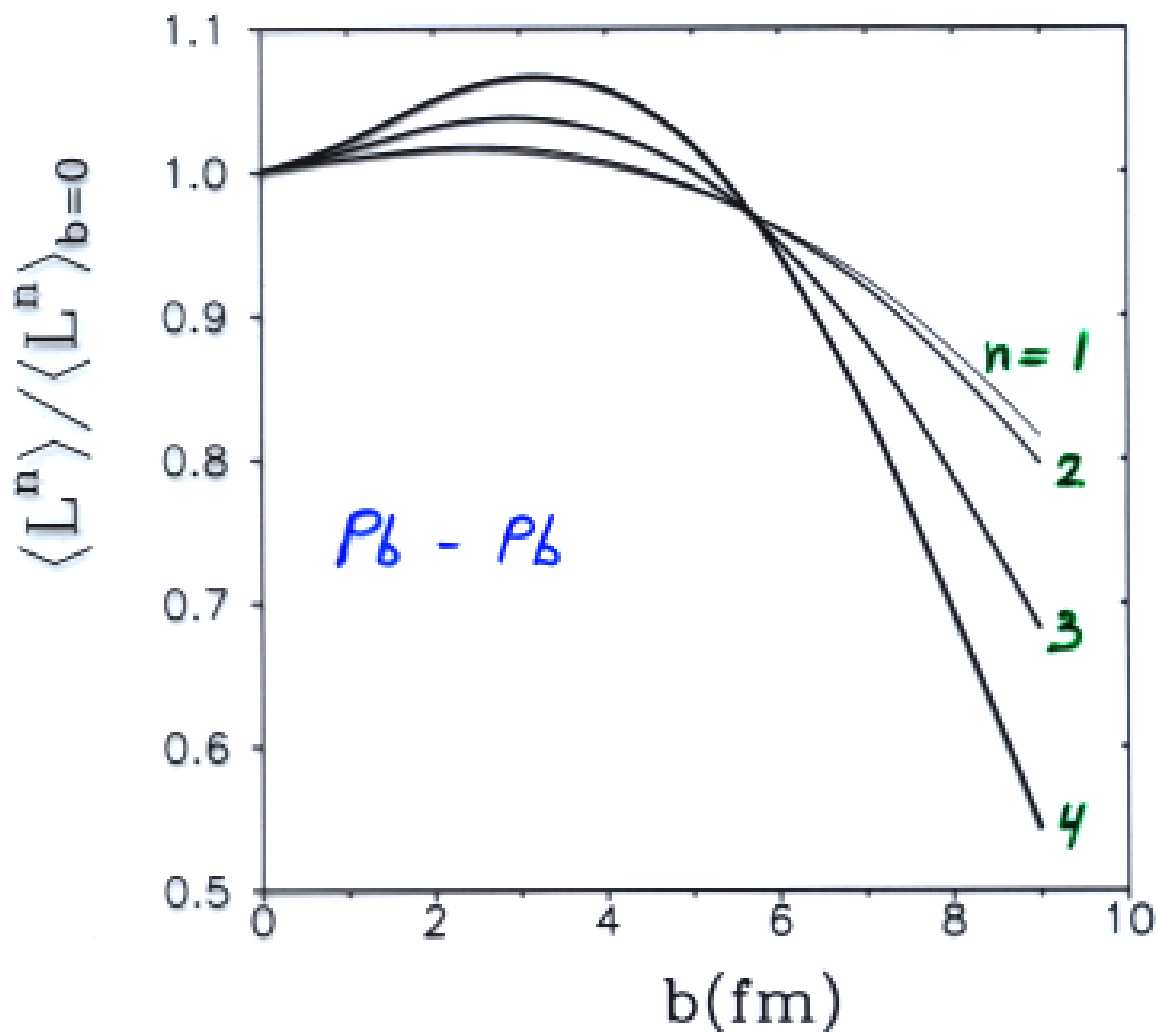
Presence of all these effects is undisputable. However, the amount of theoretical uncertainty varies and more work is needed.

Most of the effects are nonlinear and grossly enhanced for heavy nuclei:

J. Hüfner
B. K.
A. Polleri:

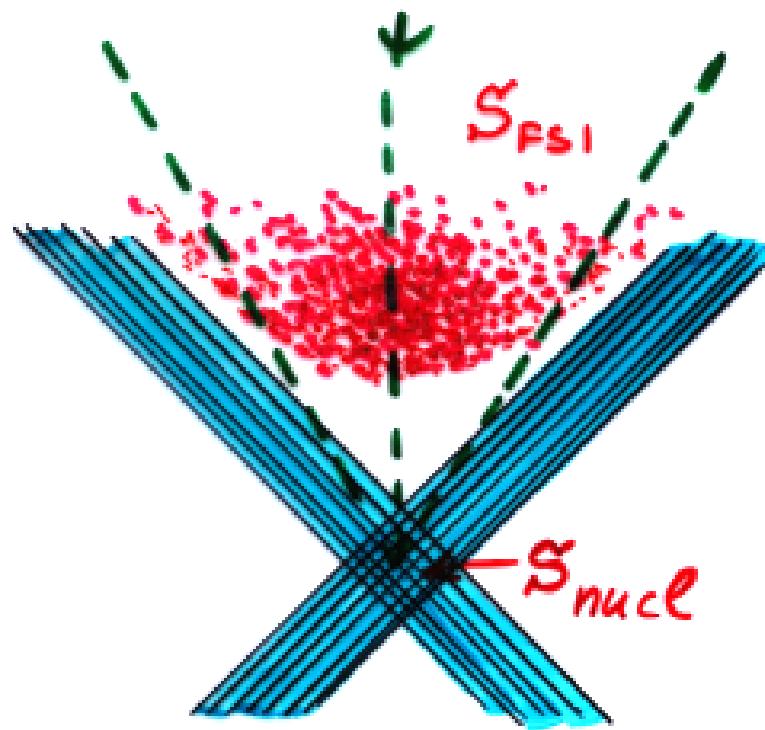


The above effects depend only on the nuclear thickness function and can provide only a flat or even rising E_T -dependence for central collisions.



The falling E_T dependence of Ψ suppression signals about interaction with the produced matter, comoving hadrons, QGP ...

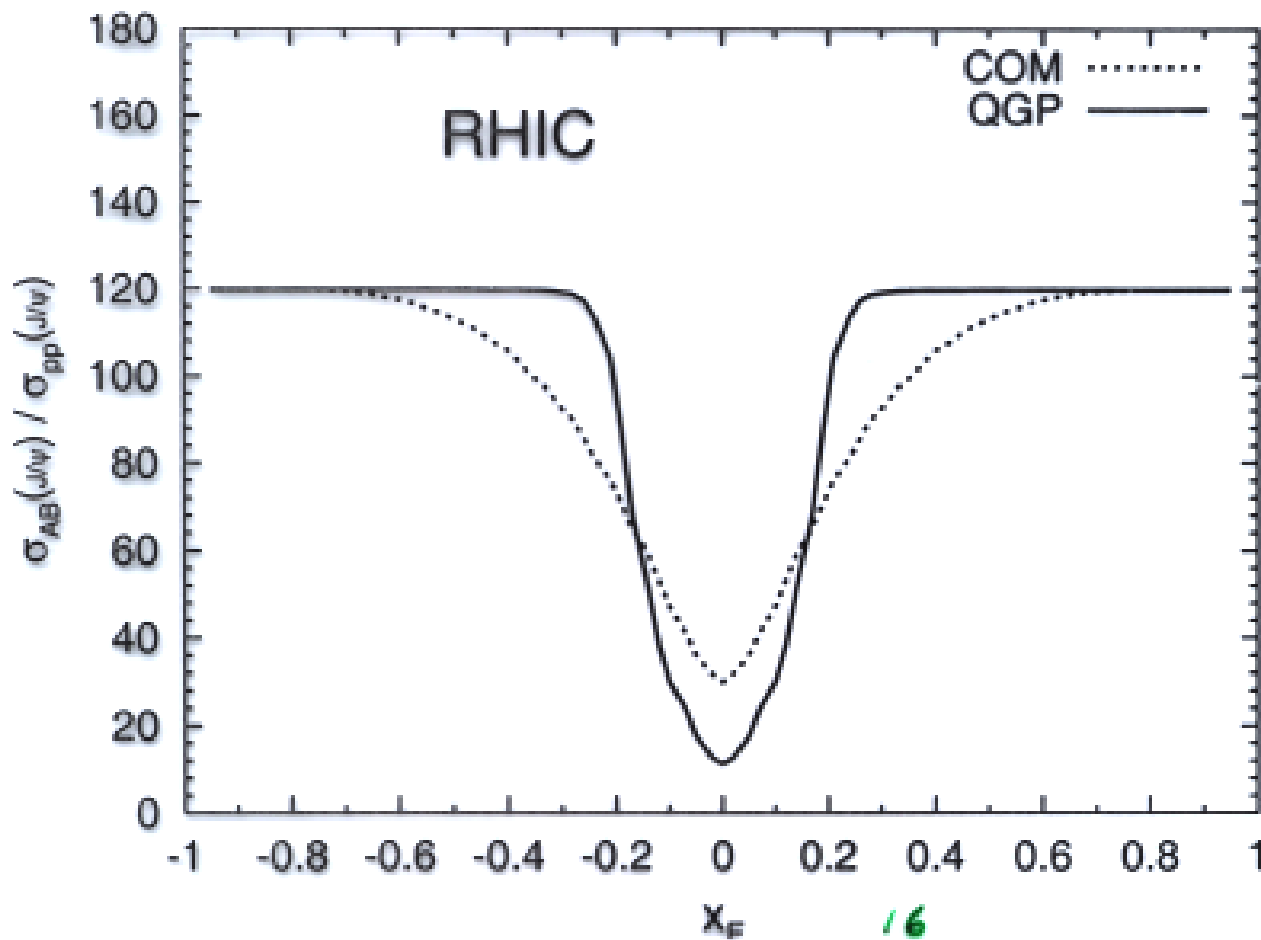
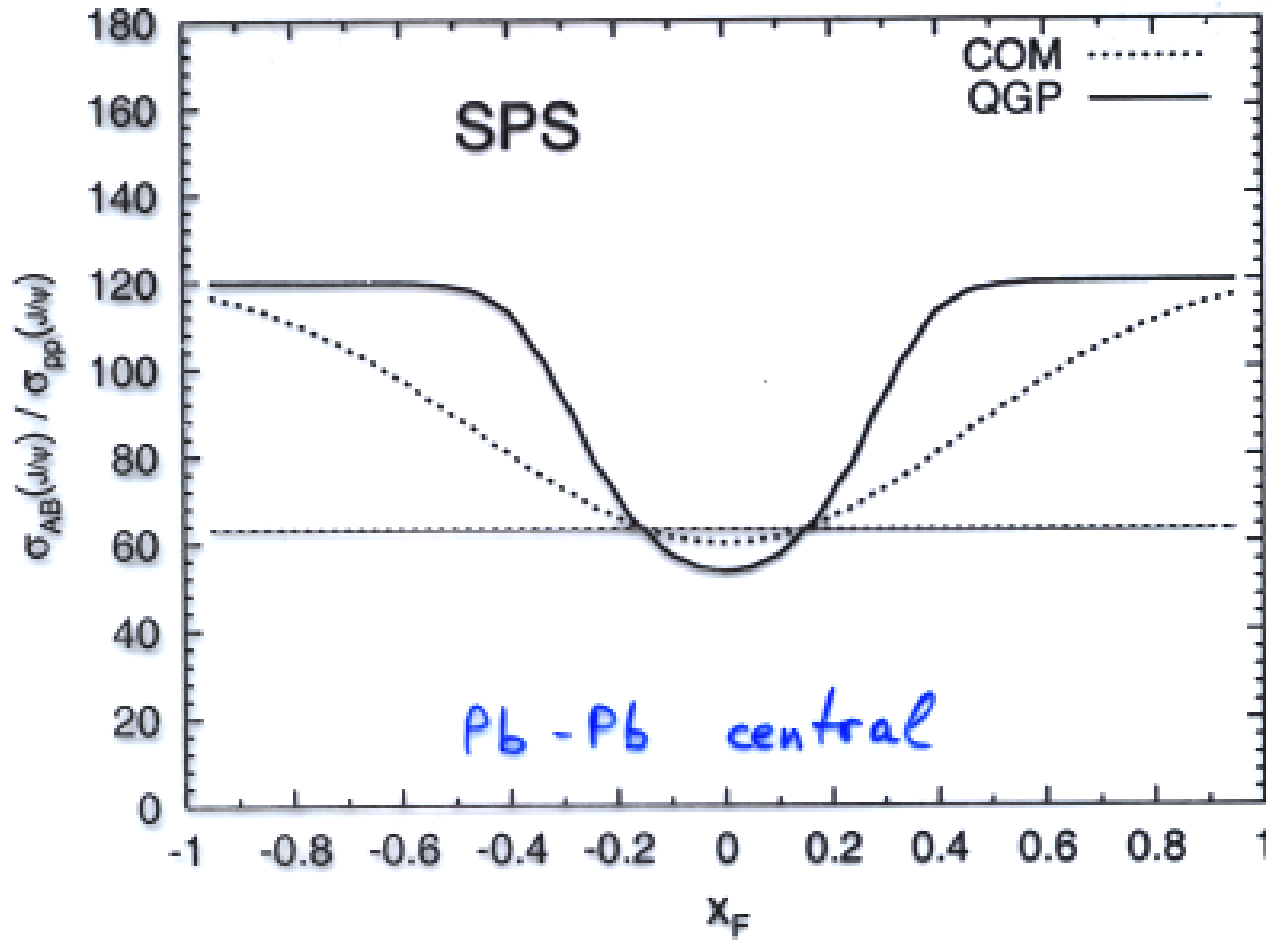
Scanning the QGP



$$S_{AB}^{\Psi} = S_{nucl} \times S_{FSI}$$

Varying x_F one scans different areas of the produced medium which is expected to be most dense at $x_F = 0$

J. Hüfner
B.K.
A. Polleri



RHIC

J. Hüfner
B. K.
A. Tarasov

Many of the effects under discussion are gone (energy loss, formation time effects, prompt gluons)

A new phenomenon becomes the main source of nuclear suppression of Ψ ,
shadowing.

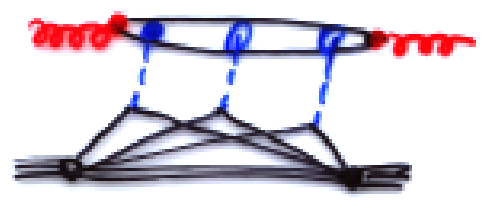
Shadowing for quarks and gluons.

Partonic interpretation is not Lorentz invariant, only observables are.

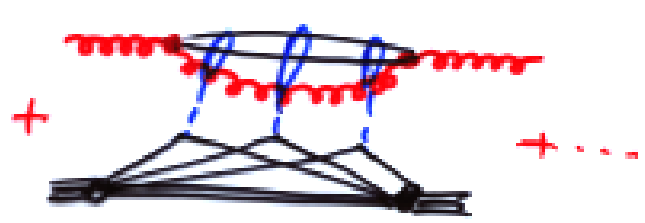
Parton model
(fusion)



Light-cone (Gribov's) formulation
(multiple interaction of Fock states)



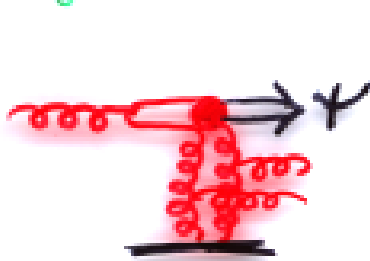
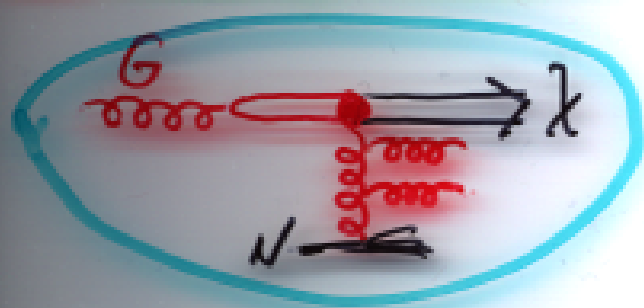
quark shadowing



gluon shadowing

(A. Mueller)

★ Hadroproduction of charmonia



still ambiguous

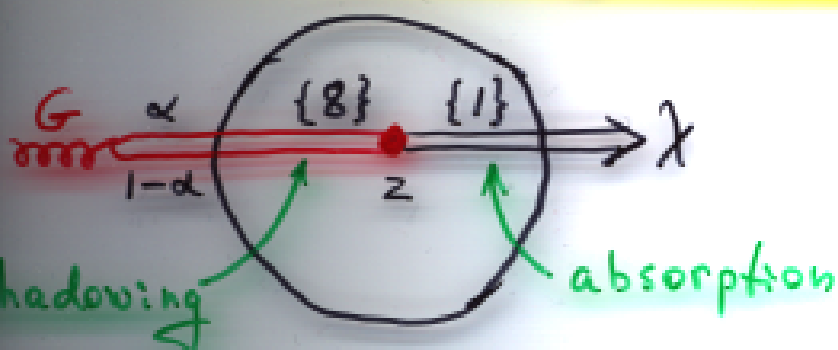
How to calculate shadowing?

$$\textcircled{1} \quad l_c = \frac{2E_\chi}{M_\chi^2} \gg R_A$$

J. Hüfner
B.K.

A. Tarasov

B.K.
A.L. Zamolodchikova
1980



$$\sigma(GA \rightarrow \chi\chi) = \int d^2b \int_{-\infty}^{\infty} dz \rho_A(b,z) |A(b,z)|^2$$

$$A(b,z) = \int_0^1 d\alpha \int d^2r_T \Phi_\chi^*(r_T, \alpha) \hat{A}(b,z, \vec{r}_T, \alpha) \Phi_G(\vec{r}_T, \alpha)$$

$$\hat{A}(b,z, \vec{r}_T, \alpha) = \exp\left[-\frac{1}{2} \sigma_{\text{abs}}(\vec{r}_T) T_+(b,z)\right] (\vec{e}_T \cdot \vec{d})$$

$$\times \exp\left[-\frac{1}{2} \sigma_{\text{shad}}(\vec{r}_T, \alpha) T_-(b,z)\right]$$

$\sigma_{abs}(r_T) = \sigma_{9\bar{9}}^N(r_T, s)$ - the universal dipole cross section
 $\propto r_T^2 (r_T \rightarrow 0)$

Well fixed by DIS data $\sigma_{9\bar{9}}^N = \sigma_0(s) (1 - e^{-\frac{r_T^2}{r_0^2(s)}})$

G-B & W
 KST

$\sigma_{shad}(r_T, \alpha) = ?$

What is the cross of $c\bar{c}$ production by a gluon?

The gluon has two Fock components

$$|G\rangle = |G\rangle_0 + |c\bar{c}\rangle$$

$c\bar{c}$ can be produced only due to difference between the amplitudes for $|G\rangle_0$ and $|c\bar{c}\rangle$

$$\sigma^{G \rightarrow c\bar{c}}(r_T, \alpha) = \left| \text{diagram}_1 - \text{diagram}_2 \right|^2 = \sum \text{diagram}_3 = \sigma_{9\bar{9}G}^N(r_T, \alpha)$$

$$\sigma_{shad}(r_T, \alpha) = \sigma_{9\bar{9}G}^N(r_T, \alpha) = \frac{9}{8} \left[\sigma_{9\bar{9}}^N(\alpha r_T) + \sigma_{9\bar{9}}^N((1-\alpha)r_T) \right] - \frac{1}{8} \sigma_{9\bar{9}}^N(r_T) \approx \frac{7}{16} \sigma(r_T)$$

$$\Phi_G(\vec{r}_T, \alpha)$$

is different from the photon wave function only by a factor

$$\Phi_G^{\mu\bar{\mu}}(\vec{r}_T, \alpha) = \frac{\sqrt{2\alpha_s}}{4\pi} \sum_c^{\mu} \hat{O} \sum_{\bar{c}}^{\bar{\mu}} K_0(m_c r_T)$$

$$\hat{O} = m_c \vec{b} \cdot \vec{e} + i(1-2\alpha)(\vec{b} \cdot \vec{n})(\vec{e} \cdot \vec{\nabla}_T) + (\vec{e} \times \vec{n}) \cdot \vec{\nabla}_T$$

$$\Phi_X(\vec{r}_T, \alpha)$$

is related by the Lorentz boost to the k wave function in the rest frame

$$\Phi_X^{\mu\bar{\mu}}(P_T, \alpha) = \sqrt{\frac{\partial P_L(P_T, \alpha)}{\partial \alpha}} \Psi_X^{\mu\bar{\mu}}(\vec{P} = \vec{P}_T + P_L \vec{n})$$

LC wave function

rest frame

$$P_L = (\alpha - \frac{1}{2}) M_{c\bar{c}}(P_T, \alpha);$$

$$M_{c\bar{c}}^2 = \frac{m_c^2 + P_T^2}{\alpha(1-\alpha)}$$

J. Küfner

Yu. Ivanov

B.K.

A. Tarasov

PR D

62(2000)094022

Important is the Melosh spin rotation

$$\text{The spin part: } S_{i,m}^{\mu\bar{\mu}} = \frac{1}{\sqrt{2}} \sum_c^{\mu} \vec{b} \cdot \vec{e}_m \sum_{\bar{c}}^{\bar{\mu}}$$

$$\sum_c^{\mu} = \hat{R}(\vec{P}_T, \alpha) \eta_c^{\mu}$$

LC \leftarrow rest frame

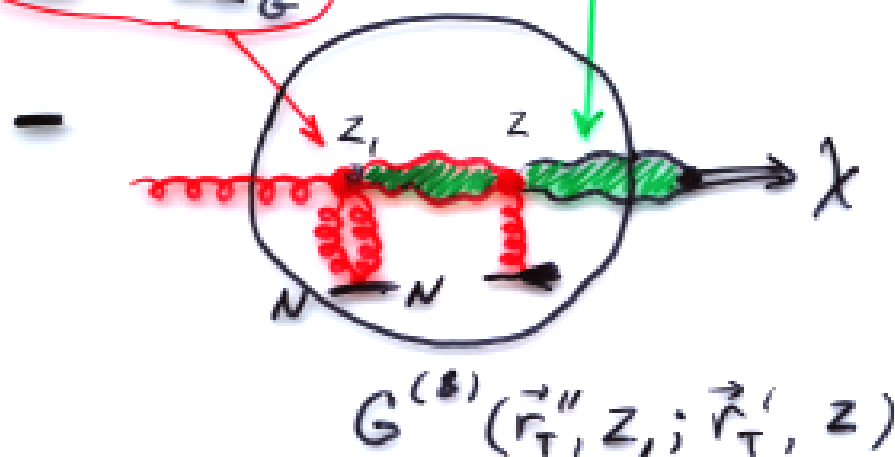
$$\sum_{\bar{c}}^{\bar{\mu}} = \hat{R}(-\vec{P}_T, 1-\alpha) \eta_{\bar{c}}^{\bar{\mu}}$$

$$\hat{R}(\vec{P}_T, \alpha) = \frac{m_c + \alpha M_{c\bar{c}} - i \vec{b} \cdot (\vec{n} + \vec{P}_T)}{\sqrt{(m_c + \alpha M_{c\bar{c}})^2 + P_T^2}}$$

General case, no restrictions for ℓ_c, ℓ_f



$q_L \approx \frac{m\chi^2}{2E_G}$



$$\therefore \frac{d}{dz} G^{(k)}(\vec{r}_T, z; \vec{r}_T', z') = \left[\frac{-\Delta_T + m_c^2}{2P_G \alpha(1-\alpha)} + V^{(k)}(\vec{r}_T, \alpha) \right] \times G^{(k)}(\vec{r}_T, z; \vec{r}_T', z')$$

$$\text{Im } V^{(k)}(\vec{r}_T, \alpha) = -\frac{1}{2} G^{(k)}(\vec{r}_T, \alpha) \rho_A(b, z)$$

$$\begin{cases} G^{(1)}(\vec{r}_T, \alpha) = G_{\text{abs}}(\vec{r}_T, \alpha) \\ G^{(2)}(\vec{r}_T, \alpha) = G_{\text{shad}}(\vec{r}_T, \alpha) \end{cases}$$

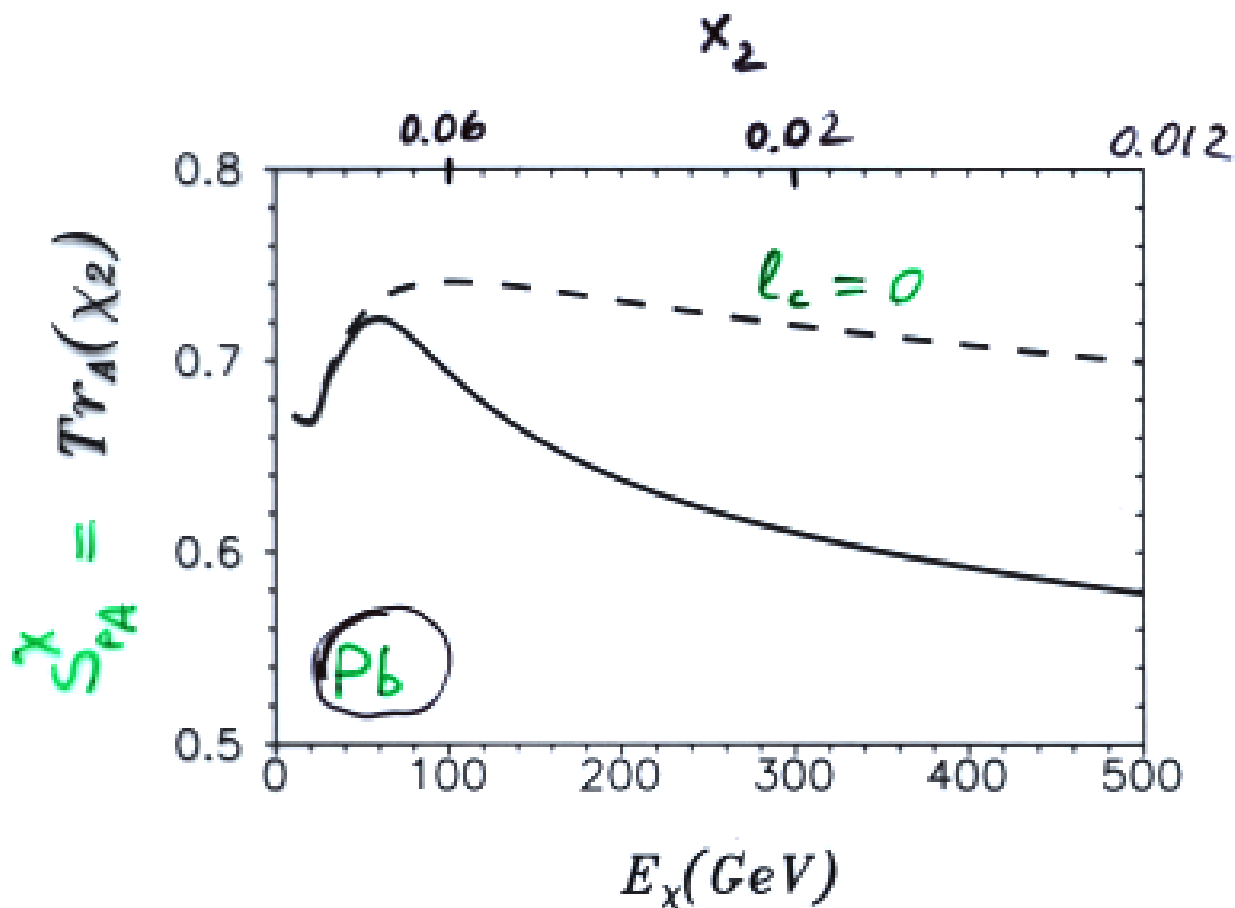
$$\text{Re } V^{(2)} = 0; \quad \text{Re } V^{(1)} = \frac{a^4(\alpha) r_T^2}{2P_G \alpha(1-\alpha)}$$

$$a^2(\alpha) = 4\alpha(1-\alpha) \omega \frac{m_c}{2}$$

Both shadowing and absorption depend only on the energy of λ

$$E_\lambda = x_1 E_P = \frac{M_X^2}{2m_N} x_2$$

x_2 scaling



One cannot borrow quark shadowing from DIS ($F_2^A(x)/F_2^N(x)$).

No QCD factorization for production of colorless states.

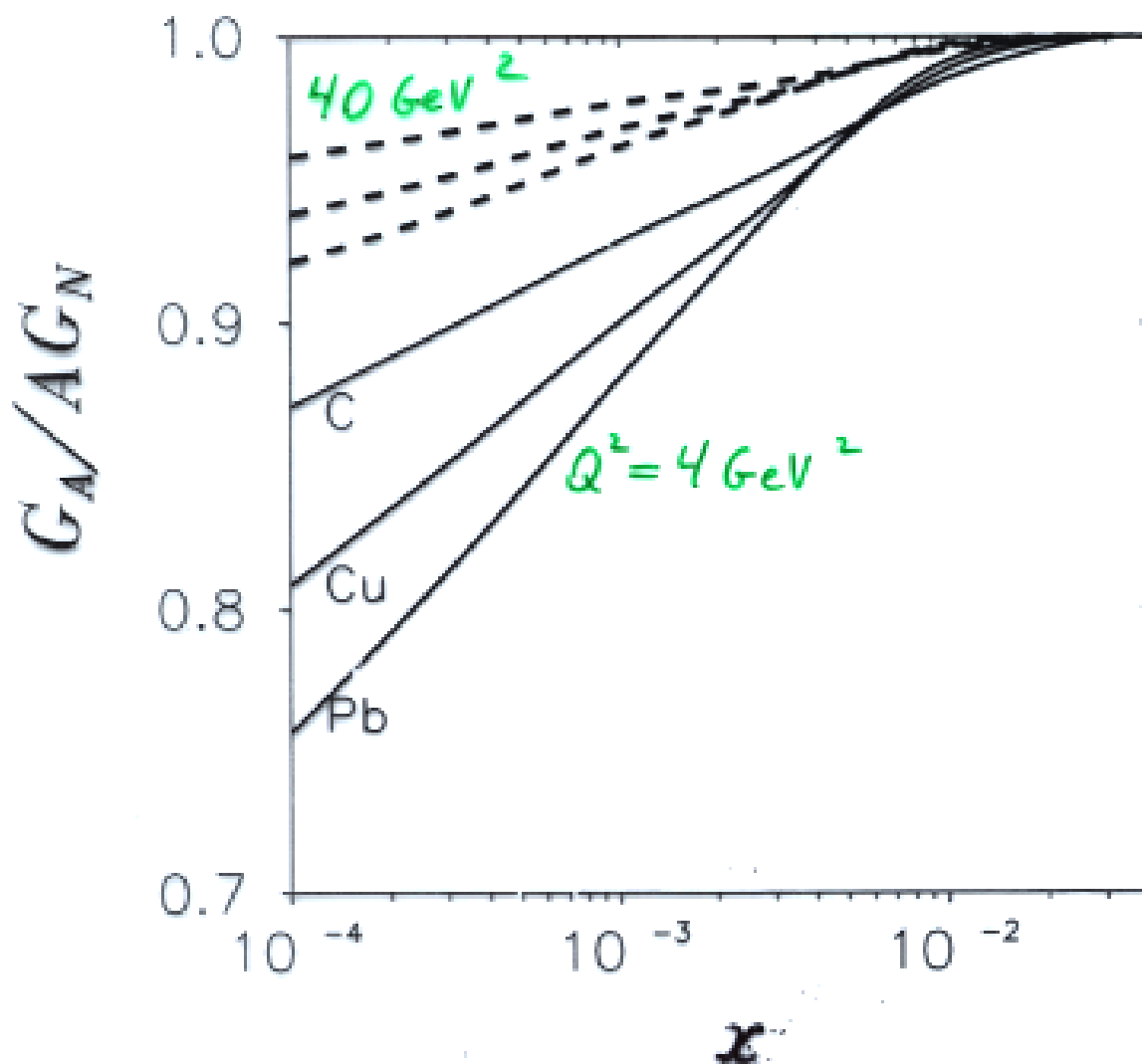
Gluon shadowing

B.K.
A. Schäfer
A. Tarasov

DIS :

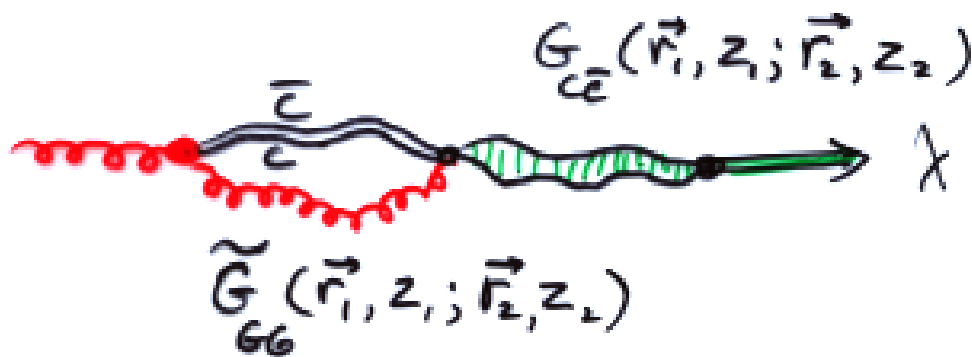


The strong nonperturbative interaction between $(q\bar{q})_s$ and G squeezes the wave packet down to the size $r_0 \sim 0.3 \text{ fm}$



Rather weak gluon shadowing in DIS

No nonperturbative gluon interaction contributes if a colorless state is produced.



The modified Green function: $\text{Re } V = 0$

The effective shadowing cross section

$$\sigma_{\text{shad}} \approx 7 \text{ mb}$$

DIS



$$\sigma_{\text{shad}} \approx 50 \text{ mb}$$

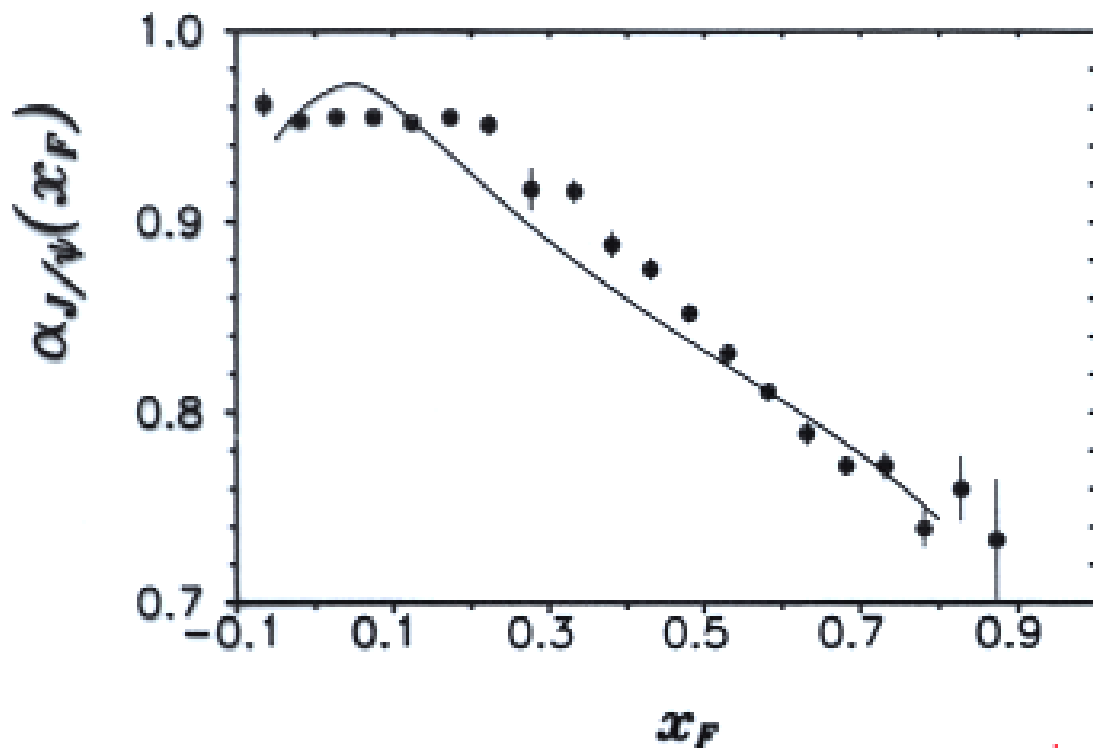
Charmonium production

Dramatic violation of the QCD scaling!

The onset of shadowing has been observed by the E772/E866 experiments at Fermilab.

Effects involved:

- shadowing, both quark and gluon
- gluon antishadowing at large x_2 (EKS)
- energy loss
- formation time effects
- energy dependence of $\sigma_{tot}^{\chi N}(s)$
- decay $\chi \rightarrow \Upsilon/\psi + \gamma$



a parameter free calculation

Again, scales in x_2

At low energies x_2 scaling is broken by energy loss.

Conclusions

- The oversimplified absorption model used at the SPS energies misses many important nonlinear effects.
- A complementary study of the x_F -dependence of charmonium suppression should provide new important information about the properties of the produced medium.
- A light cone QCD approach is developed which incorporates the effects of absorption and shadowing.
- Shadowing becomes a dominant effect at RHIC increasing nuclear suppression by a factor of few.
 x_2 scaling is predicted